

# REMOTE SENSING - RADIATIVE TRANSFER IN THE ATMOSPHERE

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FACULTY OF GEO-INFORMATION SCIENCE AND EARTH OBSERVATION



## INTRODUCTION

- Main topic: Radiative transfer
- At the end of this lesson, you can:
  - Describe the main modes of remote sensing: active and passive
  - Perform calculations of the transmission and attenuation of electromagnetic radiation
  - Explain what the concepts reflection, absorption, emission, and scattering are and where they are found in the atmosphere
  - Extract spectral information from (Earthshine) spectra
  - Perform comparative analyses of satellite reflectance vs. lab reflectance curves

## OUTLINE

- Introduction
- Recap: EMR, Blackbodies & Resolution
- Radiative transfer
  - Attenuation, transmission, reflection
  - Absorption
  - Scattering
- Types of remote sensing for EO
- Exercise

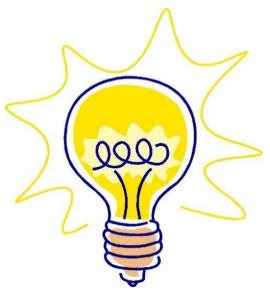
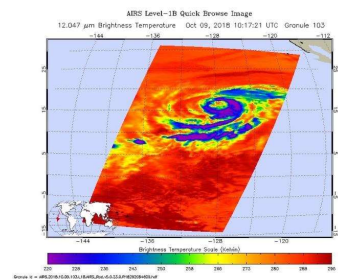
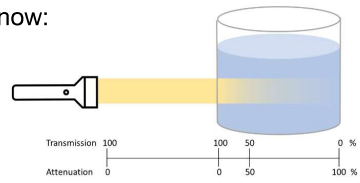


Image from ClipArtBest.com

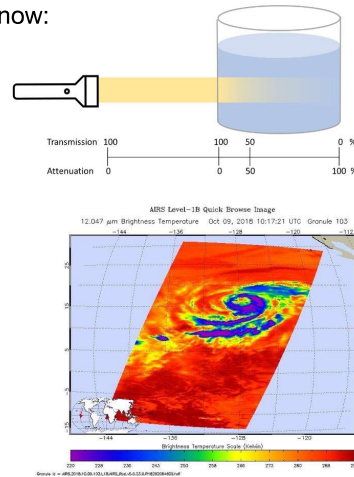
## APPLICATIONS OF RADIATIVE TRANSFER

- You need to understand RT when you want to know:
  - how far your signal will travel
  - how strong your source needs to be
  - what you're looking at (remote sensing)
  - what your radiation might do...



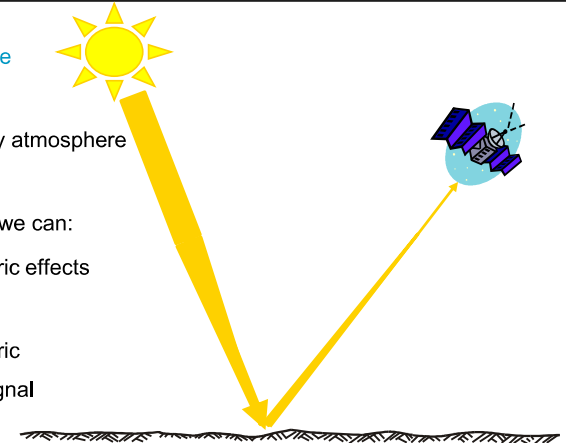
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## WHY DO WE NEED TO UNDERSTAND RT?

- Interpretation of satellite observations
- Radiation is affected by atmosphere and surface
- If we understand how, we can:
  - Remove atmospheric effects from signal
  - Retrieve atmospheric information from signal

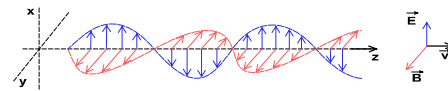


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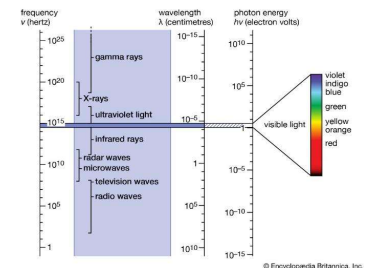
## RECAP: ELECTROMAGNETIC RADIATION

- Electromagnetic radiation consists of coupled electric and magnetic fields travelling through space at ...?
- EMR can be described as wave or particle



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- Electromagnetic radiation consists of coupled electric and magnetic fields travelling through space at ...?
- EMR can be described as wave or particle
- $c = \lambda \nu$ , therefore if  $\lambda$  is large,  $\nu$  is...?
- $Q = h\nu = \frac{hc}{\lambda}$ , therefore if  $\lambda$  is large,  $Q$  is...?

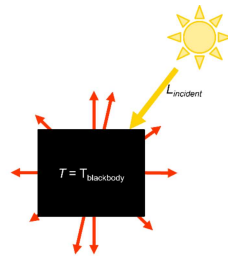


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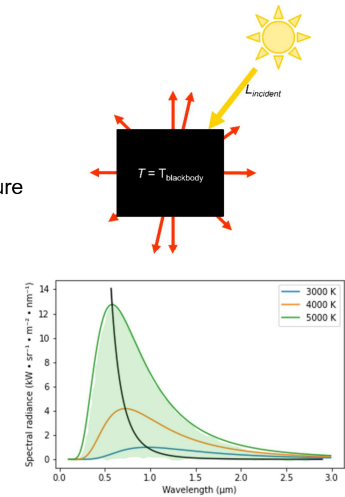
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- Every object with  $T > 0$  K emits radiation
- Blackbody:
  - Absorbs all incident radiation
  - Emitted radiation depends only on its temperature



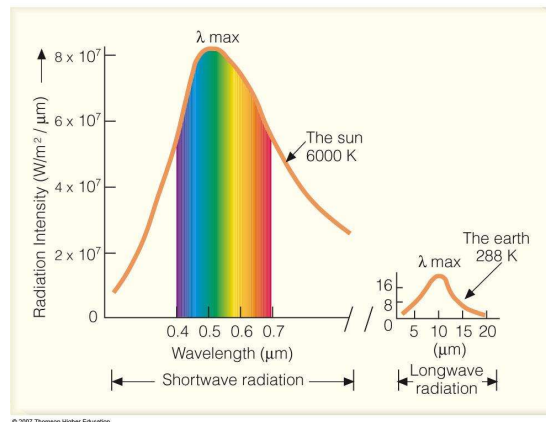
## RECAP: BLACKBODY RADIATION

- Every object with  $T > 0$  K emits radiation
- Blackbody:
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  - Emitted radiation depends only on its temperature
- Planck's law:  $L_{BB} = \frac{2hc^2}{\lambda^5} \frac{1}{e^{\frac{hc}{\lambda kT}} - 1}$
- In thermal equilibrium (no change in  $T$ ):  
 $L_{BB} = L_{incident}$
- Wien's law:  $\lambda_{max} = \frac{b}{T}$
- Stefan-Boltzmann equation:  $M = \sigma T^4$



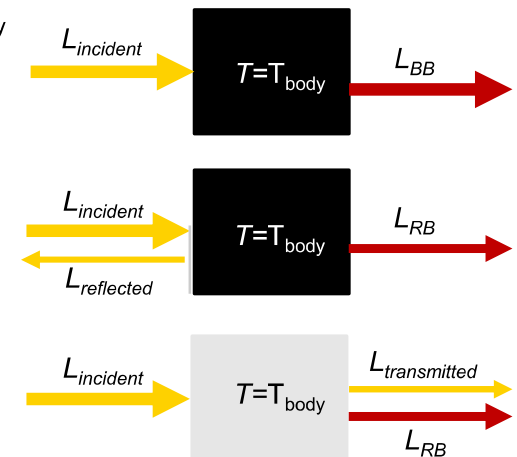
## EMISSION BY SUN AND EARTH

- Sun is much hotter than Earth ->  $\lambda_{max}$  shorter and  $M$  much greater
- Emission curves of Sun and Earth barely overlap



## REAL OBJECTS

- Real objects are not blackbodies, they
  - reflect and/or
  - transmit
- In thermal equilibrium ( $T$  constant):  
 $L_{in} = L_E + L_T + L_R$
- Emissivity  $\varepsilon_{RB} = \frac{L_{RB}}{L_{BB}}$



## QUIZ ON BLACKBODY RADIATION

How to participate?

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- 1 Go to [wooclap.com](https://wooclap.com)
- 2 Enter the event code in the top banner

Event code  
**CINZCA**



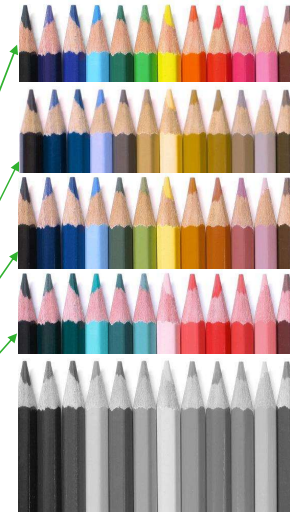
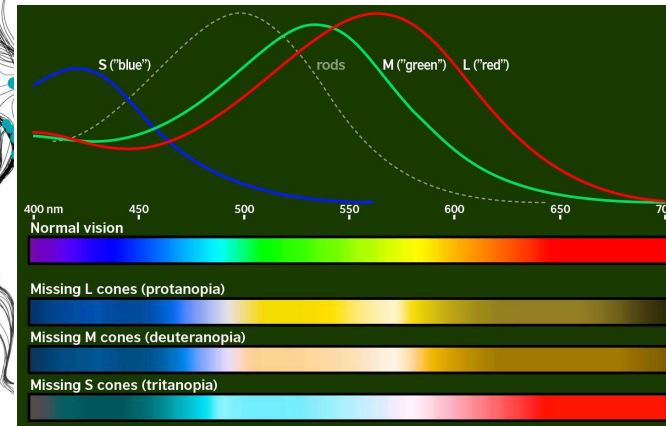
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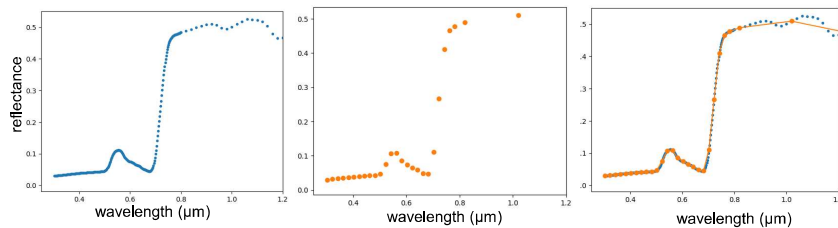
## RECAP: SPECTRAL RESOLUTION



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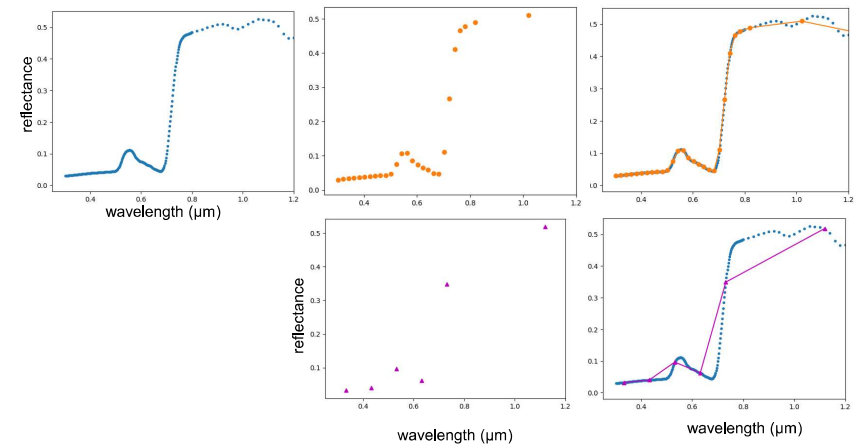
<https://factsverse.com/color-blind-people-see-world/>  
[http://www.scifun.ed.ac.uk/pages/about\\_us/shows/cb/cb-spectra-bars.jpg](http://www.scifun.ed.ac.uk/pages/about_us/shows/cb/cb-spectra-bars.jpg)

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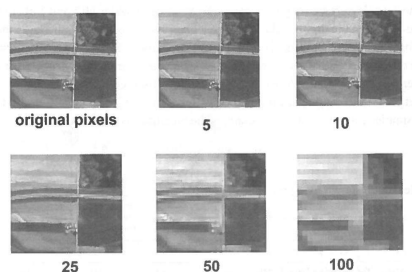


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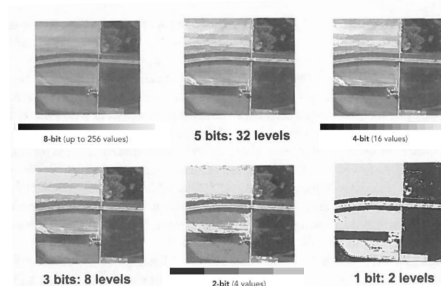


## RECAP: SPATIAL VS. RADIOMETRIC RESOLUTION

### ■ Spatial



### Radiometric



## RADIATIVE TRANSFER IN THE ATMOSPHERE

## WHAT DOES THE ATMOSPHERE DO TO RADIATION?

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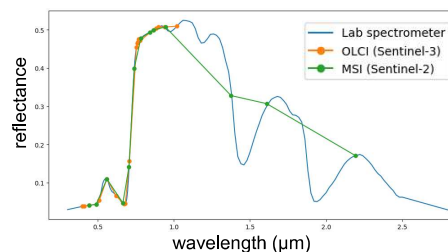
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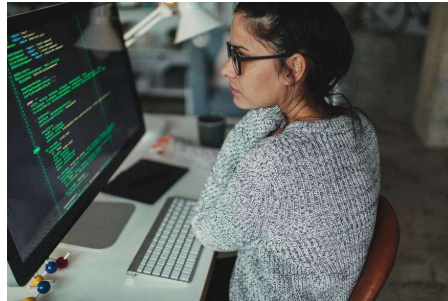
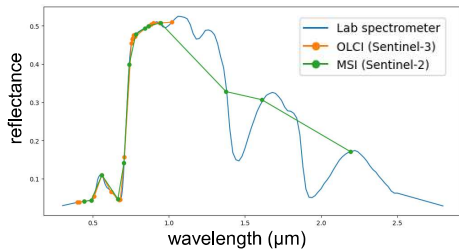
## LET ME MAKE THIS CLEAR...

- Satellites do not measure cloud cover or soil moisture content or even temperature
- Satellite instruments detect upwelling (reflected, scattered, emitted) **radiation**



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- Satellites do not measure cloud cover or soil moisture content or even temperature
- Satellite instruments detect upwelling (reflected, scattered, emitted) **radiation**
- It's up to the scientist to interpret the radiation patterns



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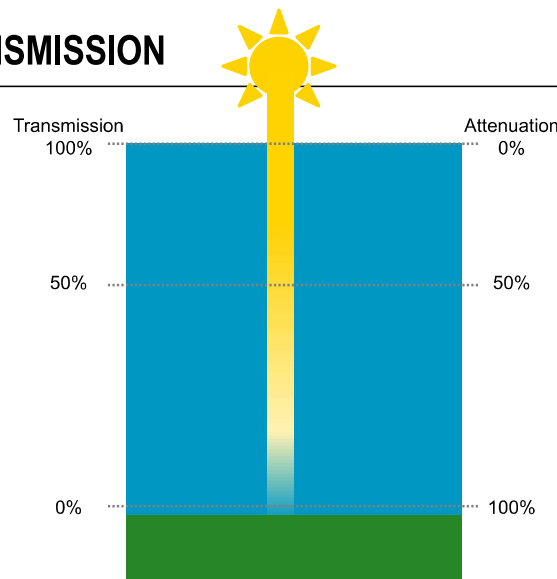
## RADIATIVE TRANSFER ATTENUATION AND TRANSMISSION



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## ATTENUATION AND TRANSMISSION

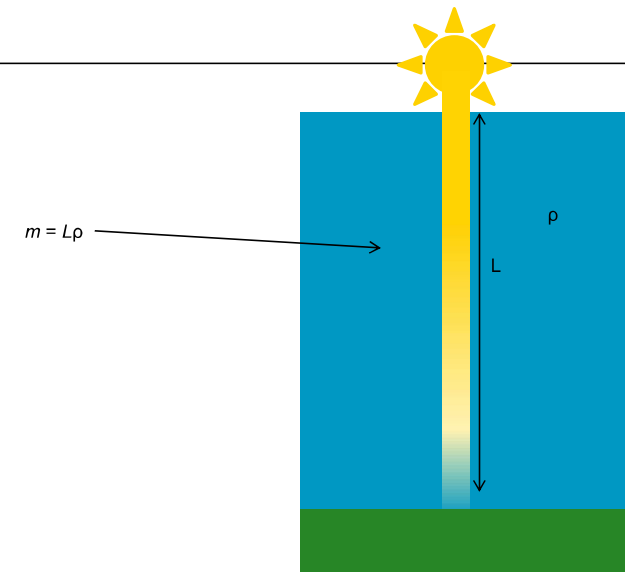
- Radiation is either **attenuated** or **transmitted** by the atmosphere
- Radiation can be attenuated by **scattering** or **absorption**



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## ATTENUATION

- Degree of attenuation (**optical air mass  $m$** ) depends on:
  - **path length** through medium ( $L$ ) and **density** of medium ( $\rho$ )
- But the atmosphere is layered...

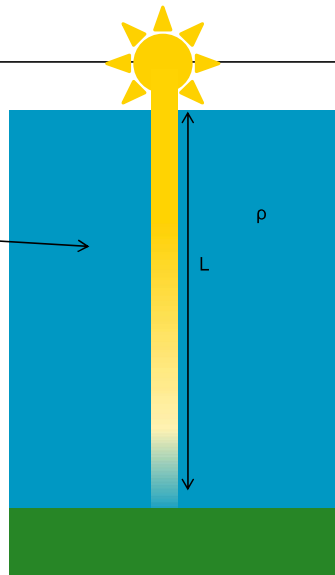


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$$m = L\rho$$



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$$m_0 = \Delta L \rho_0$$

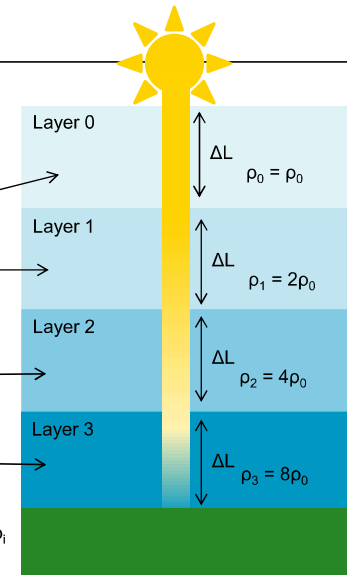
$$m_1 = \Delta L \rho_1 = 2 \Delta L \rho_0 = 2m_0$$

$$m_2 = \Delta L \rho_2 = 4 \Delta L \rho_0 = 4m_0$$

$$m_3 = \Delta L \rho_3 = 8 \Delta L \rho_0 = 8m_0$$

$$m = m_0 + m_1 + m_2 + m_3 = \sum m_i = \sum \Delta L_i \rho_i$$

$$\rightarrow m = \int \delta L \rho_i$$

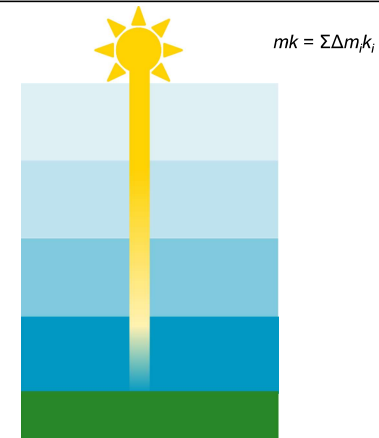


## SO HOW TO CALCULATE TRANSMISSION?

- Bouguer's law: transmission  $\tau = e^{-mk}$ 
  - $m$  = absolute air mass
  - $k$  = extinction coefficient (probability)

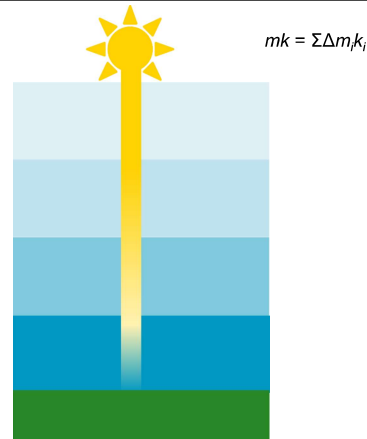
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- Hence:  $\tau = e^{-\sum_{i=0}^{i=N} m_i k_i} = \prod_{i=0}^{i=N} e^{-m_i k_i}$
- So transmissions are **multiplicative**!

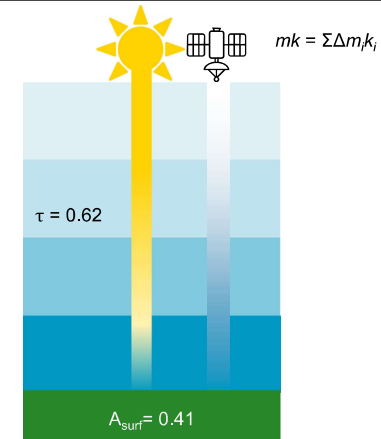


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$$\tau_{\uparrow} = e^{-\int_{z=BOA}^{z=TOA} m(z)k(z)dz}$$



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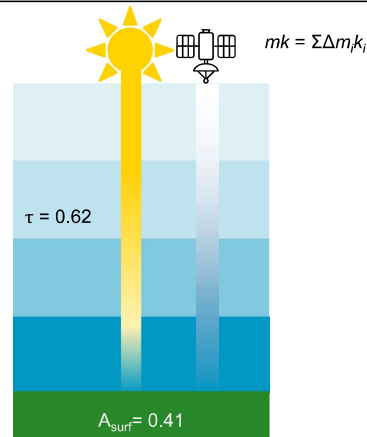
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$$\tau_{\uparrow} = e^{-\int_{z=BOA}^{z=TOA} m(z)k(z)dz}$$

$$\tau_{\downarrow} = e^{-\int_{z=TOA}^{z=BOA} m(z)k(z)dz}$$

$$\tau = \tau_{\uparrow} \cdot \tau_{\downarrow} = 0.62 \cdot 0.62 = 0.38$$

$$R = A_{surf} \tau_{\downarrow} \tau_{\uparrow} = 0.41 \cdot 0.62 \cdot 0.62 = 0.16$$

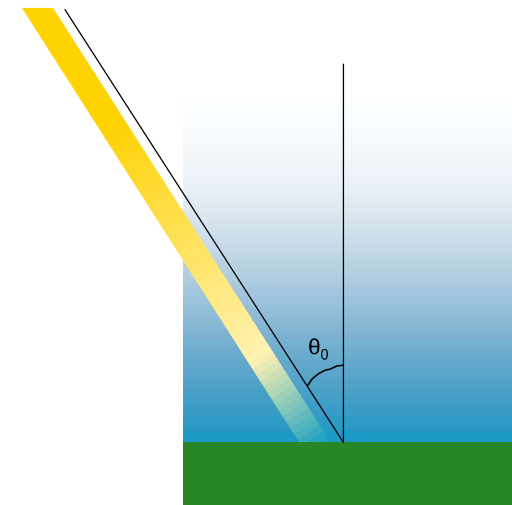


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## RELATIVE AIR MASS

- Slant paths are longer -> larger optical air mass
- Normalization using relative air mass,  $m_r$

$$m_r = \frac{L_s \cdot \rho}{L \cdot \rho} = \frac{m_s}{m}$$



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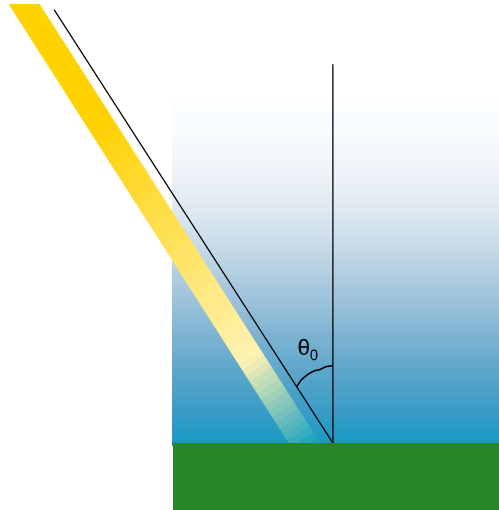


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- Radiation from Sun to Target:  $m_r = \cos \theta_0$
- So:  $m = m_r \sum \Delta L \rho_i$



## QUIZ ON ATTENUATION AND TRANSMISSION

How to participate?

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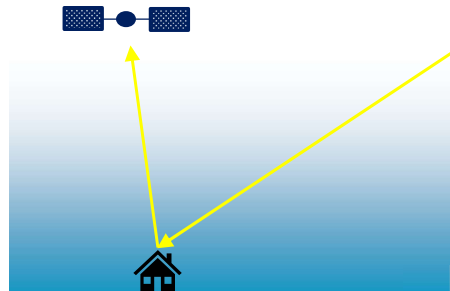
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## CALCULATE TRANSMISSIONS (1)

A medium has four constituents with optical transmissivities:

$$\tau_1 = 0.99, \tau_2 = 0.98, \tau_3 = 0.97, \tau_4 = 0.96$$

What is the total transmission of the medium?



## CALCULATE TRANSMISSIONS (2)

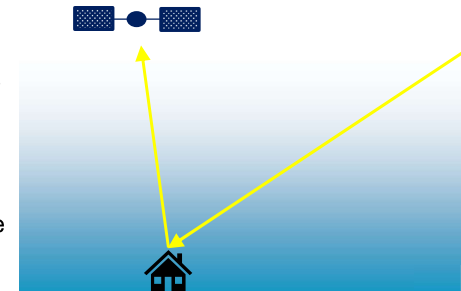
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What is the total transmission of the medium?

Bouguer's law: transmission  $\tau = e^{-mk}$

With increasing solar zenith angle, the relative air mass increases, hence the transmission...



## THAT WAS IT FOR THE EQUATIONS!

- Well, almost...

## RADIATIVE TRANSFER SCATTERING

## SCATTERING

- Interaction of radiation with large objects ( $\gg$  wavelength)
  - Reflection
  - Absorption

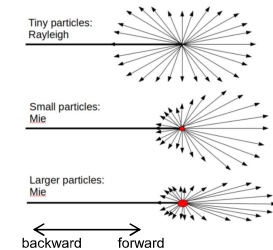


## SCATTERING

- Interaction of radiation with large objects ( $\gg$  wavelength)
  - Reflection
  - Absorption



- Interaction of radiation with small objects ( $\leq$  wavelength)
- Absorption + (partial) re-emission
- Re-emission in all directions: **scattering**



## RAYLEIGH SCATTERING

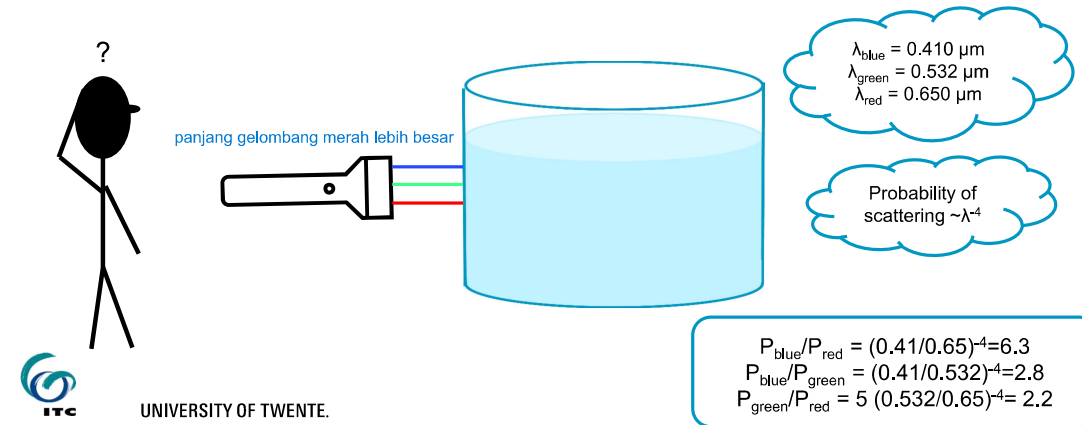
- Scattering is continuous, but depends on wavelength:
  - Size parameter  $\alpha = \pi D / \lambda$
- Rayleigh scatter:  $\alpha \ll 1$
- Probability of Rayleigh scatter  $\sim \lambda^{-4}$
- Rayleigh scatter explains why we perceive the sky as blue and sunsets as red



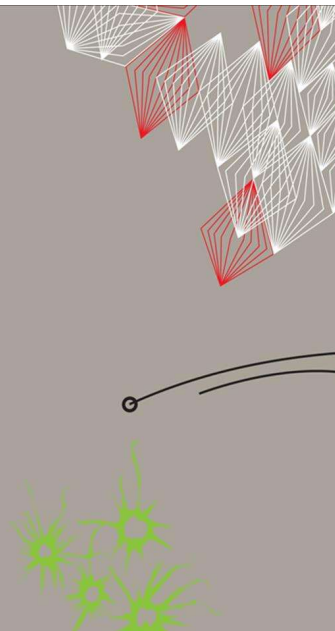
Photograph by Gabriel Parodi, ITC

## LET'S DO AN EXPERIMENT!

- What happens if we shine a blue, green or red laser into a container of milk?

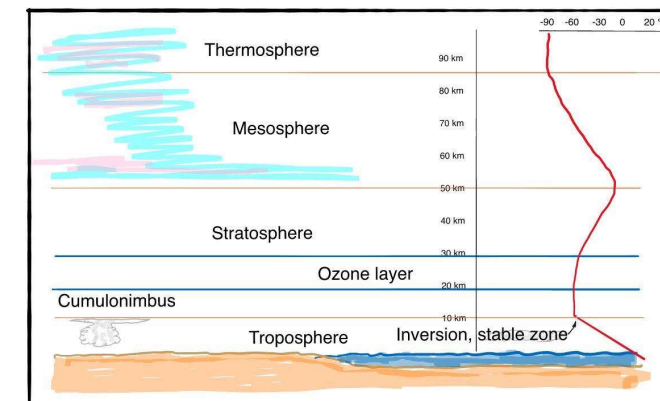


## RADIATION IN THE ATMOSPHERE ABSORPTION



## ATTENUATION IN THE ATMOSPHERE

- Atmosphere is layered:
- Layers separated by temperature profile
- Density decreases exponentially with altitude
- Most gases located in lowest layer



## ATMOSPHERIC COMPOSITION

- Nitrogen, oxygen, and argon make up 99.96% of the dry atmosphere
  - Carbon dioxide (CO<sub>2</sub>) adds 0.04%
  - Water vapour is variable, mostly in troposphere
- 
- ppm = parts per million (10<sup>6</sup>)
  - ppb = parts per billion (10<sup>9</sup>)

Constituent gas	Content % Volume or ppmv
Nitrogen (N <sub>2</sub> )	78.084 %
Oxygen (O <sub>2</sub> )	20.946 %
Argon (Ar)	0.934 %
Carbon Dioxide (CO <sub>2</sub> )	0.0416 % or 416 ppm
Other noble gases (Ne, He, Kr, Xe)	0.0024 % or 24 ppm
Hydrogen (H <sub>2</sub> )	0.00006 % or 6 ppm
Methane (CH <sub>4</sub> )	1.8 ppm
Nitrous oxide (NO <sub>2</sub> )	0.27 ppm
Ozone (O <sub>3</sub> )	0 to 0.04 ppm
Sulfur dioxide (SO <sub>2</sub> )	0.01 ppm
Nitrogen dioxide NO <sub>2</sub>	0.001 ppm
Ammonia (NH <sub>3</sub> )	0.02 ppm
Carbon Monoxide (CO)	0.09 ppm
Nitric Oxide (NO)	0.005 ppm
Hydrogen sulfide (H <sub>2</sub> S)	0.002 ppm
Nitric acid vapor	traces

Permanent  
gases

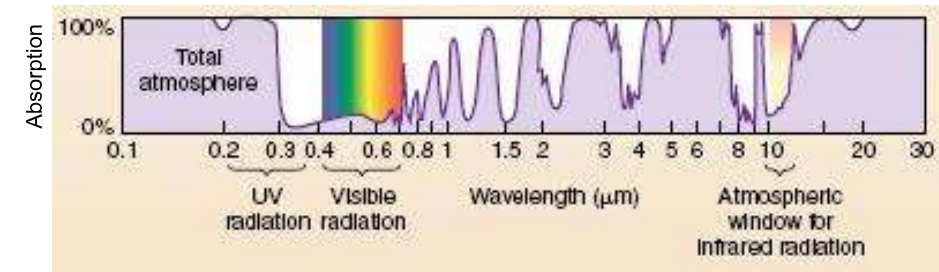
Variable  
gases



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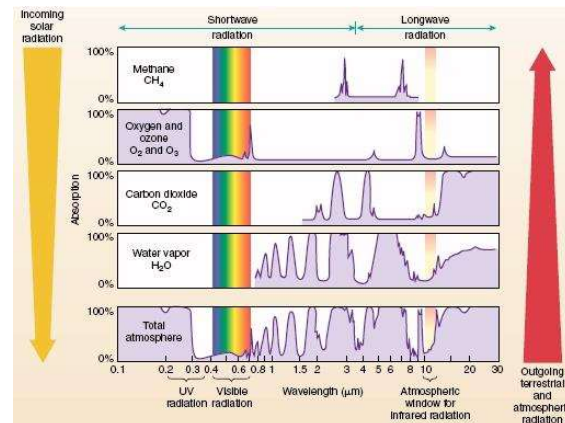
## ABSORPTION IN THE ATMOSPHERE

- Question: In which wavelength regions is the atmosphere (mostly) transparent to radiation?



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## ABSORPTION IN THE ATMOSPHERE



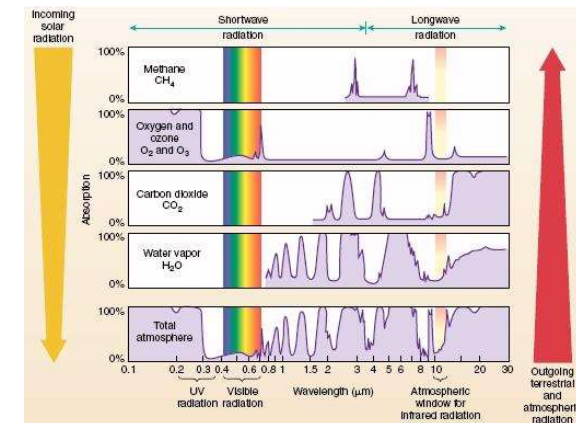
<https://cimss.ssec.wisc.edu/sage/meteorology/lesson1/AtmAbsorption.htm>



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## ABSORPTION IN THE ATMOSPHERE

- Main absorbers in the atmosphere: O<sub>3</sub>, H<sub>2</sub>O, CO<sub>2</sub>
- Most of the radiation emitted by Earth is absorbed
- In thermal equilibrium: energy received = energy emitted
- Atmospheric windows allow Earth to cool down!



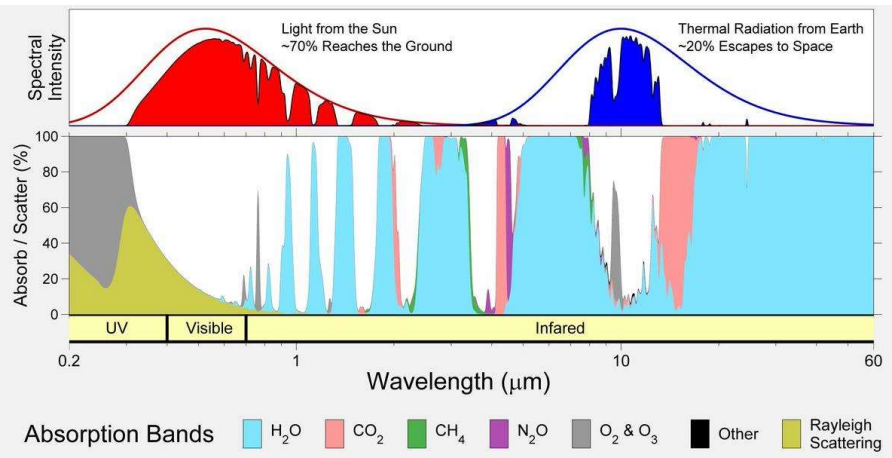
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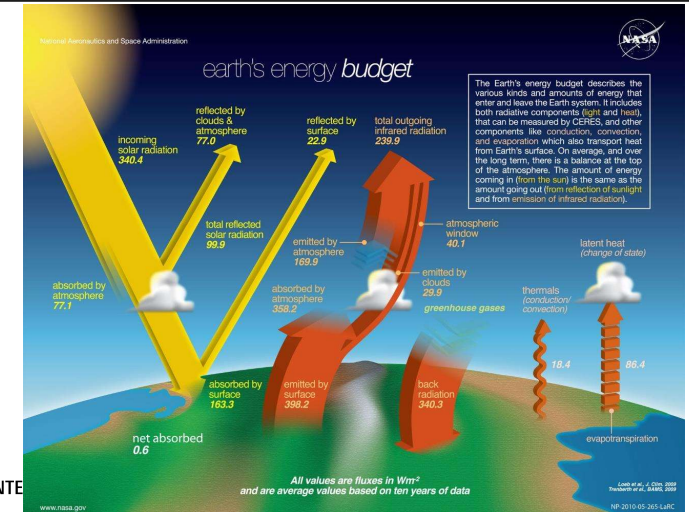
## GREENHOUSE EFFECT



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<https://twitter.com/rarohde/status/1196761516051238912>

## EXCURSION: EARTH'S ENERGY BUDGET



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## WHAT DOES THE ATMOSPHERE DO TO RADIATION?

How to participate?

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## SUMMARY

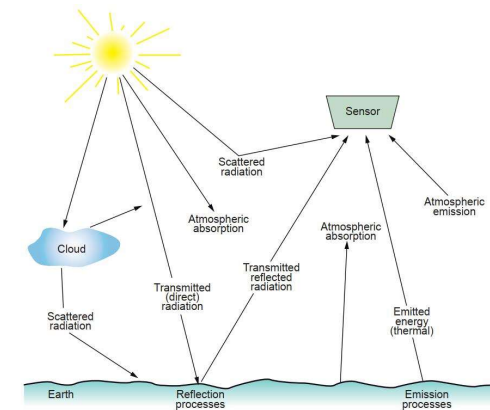


Figure 2.6  
Interactions of EM radiation  
with the atmosphere and the  
Earth's surface.

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Core book





## TODAY'S EXERCISE

---

- Supervised exercise:
  - Farzaneh Dadrass Javan
  - Srinidhi Gadde
- 10:45 - 15:30
- Rooms 2405 and 2409
- At the end of this exercise, you will be able to:
  1. Interpret reflectance spectra
  2. Find and analyse lab-based reflectance data, including metadata
  3. Extract reflectance spectra from a satellite data set using QGIS
  4. Compare spectra obtained from the lab and from satellites

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